# MOMENT-ENTROPY INEQUALITIES ${ }^{1}$ 

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#### Abstract

It is shown that the product of the Rényi entropies of two independent random vectors provides a sharp lower bound for the expected value of the moments of the inner product of the random vectors. This new inequality contains important geometry (such as extensions of one of the fundamental affine isoperimetric inequalities, the Blaschke-Santaló inequality).


1. Introduction. Vitale (1996a, b, 2001) presents evidence of a surprising connection between probability and analytic convex geometry. In this paper we contribute additional evidence of this unexpected link between these subjects.

It will be shown that the product of the Rényi entropies of two independent random vectors provides a sharp lower bound for the expected value of the moments of the inner product of the random vectors. Our new inequality encodes important geometry. For example, a nontechnical version of the Blaschke-Santaló inequality for compact sets is but one special case.

This paper deals with random vectors in Euclidean $n$-space, $\mathbb{R}^{n}$. For vectors $x, y \in \mathbb{R}^{n}$, let $x \cdot y$ denote their inner product. If $\phi \in G L(n)$, then we use $\phi^{-t}$ to denote the inverse of the transpose of $\phi$.

If $X$ is a random vector in $\mathbb{R}^{n}$ with density $f$, then for $\lambda>0$, the $\lambda$-Rényi entropy of $X$ is defined [see, e.g., Cover and Thomas (1992) and Gardner (2002)] by

$$
\operatorname{Ent}_{\lambda}(X)= \begin{cases}\frac{1}{1-\lambda} \log \int_{\mathbb{R}^{n}} f(x)^{\lambda} d x, & \lambda \neq 1 \\ -\int_{\mathbb{R}^{n}} f(x) \log f(x) d x, & \lambda=1\end{cases}
$$

The $\lambda$-Rényi entropy power of $X$ is defined as

$$
N_{\lambda}(X)=e^{\operatorname{Ent}_{\lambda}(X)} \quad \text { with } \quad N_{\infty}(X)=\lim _{\lambda \rightarrow \infty} N_{\lambda}(X)
$$

A random vector in $\mathbb{R}^{n}$, with density function $f$, has finite pth moment provided that

$$
\int_{\mathbb{R}^{n}}|x|^{p} f(x) d x<\infty
$$

where $|x|$ denotes the ordinary Euclidean norm of $x \in \mathbb{R}^{n}$.

[^0]For $\phi \in G L(n)$, and $p, \lambda>0$, define densities $\phi_{p, \lambda}: \mathbb{R}^{n} \rightarrow \mathbb{R}$ by

$$
\phi_{p, \lambda}(x)= \begin{cases}b\left(1+|\phi x|^{p}\right)^{1 /(\lambda-1)}, & \lambda<1 \\ b e^{-|\phi x|^{p}}, & \lambda=1 \\ b\left(1-|\phi x|^{p}\right)_{+}^{1 /(\lambda-1)}, & \lambda>1\end{cases}
$$

where $(z)_{+}=\max \{0, z\}$, and in each case $b=b_{p, \lambda}$ is chosen so that $\phi_{p, \lambda}$ is a density.

We shall prove an extended version of the following:
THEOREM. Suppose real $p \geq 1$ and real $\lambda>\frac{n}{n+p}$ are fixed. If $X$ and $Y$ are independent random vectors in $\mathbb{R}^{n}$ that have finite pth moment, then

$$
E\left(|X \cdot Y|^{p}\right) \geq c_{1}\left[N_{\lambda}(X) N_{\lambda}(Y)\right]^{p / n}
$$

where the best possible $c_{1}$ is given by

$$
\begin{aligned}
c_{1}= & \frac{2}{n \pi^{p+1 / 2}} \Gamma\left(\frac{p+1}{2}\right) \Gamma\left(\frac{n+p}{2}\right)^{-1} \Gamma\left(\frac{n}{2}+1\right)^{1+2 p / n} c_{0}^{2}, \\
c_{0}^{-n / p}= & \begin{array}{ll}
\frac{n}{p}\left(1-\frac{n(1-\lambda)}{p \lambda}\right)^{1 /(\lambda-1)} & \\
\quad \times\left(\frac{p \lambda}{n(1-\lambda)}-1\right)^{n / p} B\left(\frac{n}{p}, \frac{1}{1-\lambda}-\frac{n}{p}\right), & \lambda<1, \\
\left(\frac{p e}{n}\right)^{n / p} \Gamma\left(\frac{n}{p}+1\right), & \lambda=1, \\
\frac{n}{p}\left(1+\frac{n(\lambda-1)}{p \lambda}\right)^{1 /(\lambda-1)} & \\
\quad \times\left(\frac{p \lambda}{n(\lambda-1)}+1\right)^{n / p} B\left(\frac{n}{p}, \frac{\lambda}{\lambda-1}\right), & \lambda>1 .
\end{array}
\end{aligned}
$$

Equality occurs if and only if $X$ has density a.e. $\phi_{p, \lambda}$ and $Y$ has density a.e. $\left(a \phi^{-t}\right)_{p, \lambda}$, with $\phi \in G L(n)$ as $a>0$.

If $K \subset \mathbb{R}^{n}$ is a compact set with volume (i.e., Lebesgue measure) $V(K)$ and $X$ has density $\mathbf{1}_{K} / V(K)$, then trivially $N_{\infty}(X)=V(K)$. Now if $K, L \subset \mathbb{R}^{n}$ are compact and we let $X, Y$ have densities $\mathbf{1}_{K} / V(K), \mathbf{1}_{L} / V(L)$, then letting $\lambda \rightarrow \infty$ and $p \rightarrow \infty$ in the theorem gives the following inequality.

Corollary. If $K, L \subset \mathbb{R}^{n}$ are compact, then

$$
\omega_{n}^{2} \max _{x \in K, y \in L}|x \cdot y|^{n} \geq V(K) V(L)
$$

Here $\omega_{n}=\pi^{n / 2} / \Gamma(1+n / 2)$ denotes the volume of the unit ball in $\mathbb{R}^{n}$. If $K$ is an origin-symmetric convex body and $L$ is the polar of $K$, then the above inequality is the classical Blaschke-Santaló inequality, with sharp constant. See, for example, the books of Gardner (1995), Leichtweiß (1998), Schneider (1993) and Thompson (1996) [and also the article of Bourgain and Milman (1987)] for references regarding the Blaschke-Santaló inequality.
2. Dual mixed volumes of random vectors. Each nonnegative $\rho \in L_{q}\left(S^{n-1}\right)$ defines a star-shaped set $K=K_{\rho} \subset \mathbb{R}^{n}$ by

$$
K=\left\{r u: 0 \leq r \leq \rho(u) \text { with } u \in S^{n-1}\right\} .
$$

The set $K$ is called the $L_{q}$-star generated by $\rho$ and the function $\rho$ is called the radial function of $K$ (and is often written as $\rho_{K}$ to indicate its relationship to $K$ ). We will not distinguish between $L_{q}$-stars whose radial functions are a.e. equal.

It is convenient to extend the definition of the radial function from $S^{n-1}$ to $\mathbb{R}^{n} \backslash\{0\}$ by making it homogeneous of degree -1 . Thus, for an $L_{q}$-star $K \subset \mathbb{R}^{n}$, the radial function $\rho_{K}: \mathbb{R}^{n} \backslash\{0\} \rightarrow[0, \infty)$ can be defined by $\rho_{K}(x)=\max \{r \geq 0$ : $r x \in K\}$. From this it follows immediately that if $\phi \in G L(n)$, then the radial function of the star $\phi K=\{\phi x: x \in K\}$ is given by $\rho_{\phi K}(x)=\rho_{K}\left(\phi^{-1} x\right)$ for all $x \neq 0$. From this, and the homogeneity (of degree -1 ) of the radial function, it follows immediately that $E$ is an origin-centered ellipsoid of positive volume if and only if there exists a $\phi \in G L(n)$ such that $1 / \rho_{E}(x)=|\phi x|$, for all $x \neq 0$.

Elements of the dual Brunn-Minkowski theory of $L_{q}$-stars were studied by Klain (1996, 1997). Other extensions for the dual Brunn-Minkowski theory have been considered by Gardner and Volčič (1994) and Gardner, Vedel Jensen and Volčič (2003). In particular, the latter paper defines dual mixed volumes of bounded Borel sets in terms of moments.

An $L_{q}$-star whose radial function is both positive and continuous is called a star body. A star body that is convex is called a convex body. Note that throughout, convex bodies are assumed to contain the origin in their interiors.

If $K$ is a convex body in $\mathbb{R}^{n}$, then the polar, $K^{*}$, of $K$ is defined by

$$
K^{*}=\left\{x \in \mathbb{R}^{n}: x \cdot y \leq 1 \text { for all } y \in K\right\}
$$

It follows immediately from this definition that if $\phi \in G L(n)$, then $(\phi K)^{*}=$ $\phi^{-t} K^{*}$. From this we see that $\rho_{1}, \rho_{2}$ are radial functions of polar reciprocal origincentered ellipsoids if and only if there exists a $\phi \in G L(n)$ such that

$$
\begin{align*}
& 1 / \rho_{1}(x)=|\phi x| \quad \text { and } \\
& 1 / \rho_{2}(x)=\left|\phi^{-t} x\right| \tag{2.1}
\end{align*}
$$

for all $x \neq 0$.

Suppose $p>0$. Define the dual mixed volume $\tilde{V}_{-p}(K, L)$ of an $L_{n+p}$-star $K$ and a star body $L$ by

$$
\begin{align*}
\tilde{V}_{-p}(K, L) & =\frac{1}{n} \int_{S^{n-1}} \rho_{K}(u)^{n+p} \rho_{L}(u)^{-p} d u  \tag{2.2}\\
& =\frac{n+p}{n} \int_{K} \rho_{L}(x)^{-p} d x
\end{align*}
$$

where in the first integral the integration is with respect to Lebesgue measure on $S^{n-1}$ and in the second integral the integration is with respect to Lebesgue measure on $\mathbb{R}^{n}$. Obviously, for each star body $L$,

$$
\begin{equation*}
\tilde{V}_{-p}(L, L)=V(L) \tag{2.3}
\end{equation*}
$$

From (2.2) and the Hölder inequality we see that if $K$ is an $L_{n+p}$-star and $L$ is a star body, then we have the dual mixed volume inequality

$$
\begin{equation*}
\tilde{V}_{-p}(K, L)^{n} \geq V(K)^{n+p} V(L)^{-p} \tag{2.4}
\end{equation*}
$$

with equality if and only if there exists a $c>0$ such that, a.e., $\rho_{K}=c \rho_{L}$. Thus, for the volume of each $L_{n+p}$-star, $K$, we have

$$
\begin{equation*}
V(K)^{(n+p) / n}=\operatorname{Inf}\left\{\tilde{V}_{-p}(K, Q): Q \text { is a star body with } V(Q)=1\right\} \tag{2.5}
\end{equation*}
$$

If $X$ is a random vector in $\mathbb{R}^{n}$, that has finite $p$ th moment and $L$ is star body in $\mathbb{R}^{n}$, then define the dual mixed volume $\tilde{V}_{-p}(X, L)$ by

$$
\begin{equation*}
\tilde{V}_{-p}(X, L)=\frac{n+p}{n} \int_{\mathbb{R}^{n}} \rho_{L}(x)^{-p} f(x) d x \tag{2.6}
\end{equation*}
$$

where $f$ is the density function $X$. It will be on occasion convenient to write $\tilde{V}_{-p}(f, L)$ rather than $\tilde{V}_{-p}(X, L)$.

For $p, \lambda>0$, define $p_{\lambda}: \mathbb{R}^{+} \rightarrow \mathbb{R}^{+}$by

$$
p_{\lambda}(s)= \begin{cases}\left(1+s^{p}\right)^{1 /(\lambda-1)}, & \lambda<1, \\ e^{-s^{p}}, & \lambda=1, \\ \left(1-s^{p}\right)_{+}^{1 /(\lambda-1)}, & \lambda>1 .\end{cases}
$$

We shall use the following lemma:

Lemma 2.1. Suppose $K$ is a star body in $\mathbb{R}^{n}$. For real $a, p, \lambda>0$, with $\lambda>n /(n+p)$,

$$
\int_{\mathbb{R}^{n}} \rho_{K}^{-p}(x) p_{\lambda}\left(a / \rho_{K}(x)\right) d x=a^{-(n+p)} \alpha_{1} V(K)
$$

where

$$
\begin{align*}
\alpha_{1} & =n \int_{0}^{\infty} s^{n+p-1} p_{\lambda}(s) d s \\
& = \begin{cases}\frac{n}{p} B\left(\frac{n+p}{p}, \frac{\lambda}{1-\lambda}-\frac{n}{p}\right), & \lambda<1, \\
\frac{n}{p} \Gamma\left(\frac{n+p}{p}\right), & \lambda=1, \\
\frac{n}{p} B\left(\frac{n+p}{p}, \frac{\lambda}{\lambda-1}\right), & \lambda>1 .\end{cases} \tag{2.7}
\end{align*}
$$

Proof. Rewrite the integral over $\mathbb{R}^{n}$ as an integral over $S^{n-1} \times(0, \infty)$,

$$
\begin{aligned}
& \int_{\mathbb{R}^{n}} \rho_{K}^{-p}(x) p_{\lambda}\left(a / \rho_{K}(x)\right) d x \\
& \quad=\int_{S^{n-1}} \int_{0}^{\infty} \rho_{K}^{-p}(r u) p_{\lambda}\left(a / \rho_{K}(r u)\right) r^{n-1} d r d u
\end{aligned}
$$

and observe that the inner integral is easily evaluated. Specifically, for fixed $u \in S^{n-1}$, make the change of variable $s=\left[a / \rho_{K}(u)\right] r=a / \rho_{K}(r u)$ and observe

$$
\begin{aligned}
\int_{0}^{\infty} & \rho_{K}^{-p}(r u) p_{\lambda}\left(a / \rho_{K}(r u)\right) r^{n-1} d r \\
& =a^{-(n+p)} \rho_{K}^{n}(u) \int_{0}^{\infty} s^{n+p-1} p_{\lambda}(s) d s=\alpha_{1} a^{-(n+p)} \frac{1}{n} \rho_{K}^{n}(u)
\end{aligned}
$$

Thus, if $K$ is a star body and $b$ is chosen so that $b p_{\lambda}\left(a / \rho_{K}\right)$ is a probability density and $a>0$, then

$$
\tilde{V}_{-p}\left(b p_{\lambda}\left(a / \rho_{K}\right), K\right)=(1+p / n) b a^{-(n+p)} \alpha_{1} V(K) .
$$

We shall need the following:
Lemma 2.2. Suppose $K$ is a star body in $\mathbb{R}^{n}$. For $a, p, \lambda>0$, with $\lambda>n /(n+p)$,

$$
\int_{\mathbb{R}^{n}} p_{\lambda}\left(a / \rho_{K}(x)\right) d x=a^{-n} \alpha_{2} V(K)
$$

where

$$
\alpha_{2}=n \int_{0}^{\infty} s^{n-1} p_{\lambda}(s) d s= \begin{cases}\frac{n}{p} B\left(\frac{n}{p}, \frac{1}{1-\lambda}-\frac{n}{p}\right), & \lambda<1,  \tag{2.8}\\ \frac{n}{p} \Gamma\left(\frac{n}{p}\right), & \lambda=1, \\ \frac{n}{p} B\left(\frac{n}{p}, \frac{\lambda}{\lambda-1}\right), & \lambda>1 .\end{cases}
$$

The proof is similar to that of Lemma 2.1.
We shall also use the following lemma:
Lemma 2.3. Suppose $K$ is a star body in $\mathbb{R}^{n}$. For $a, p, \lambda>0$, with $\lambda>n /(n+p)$,

$$
\int_{\mathbb{R}^{n}} p_{\lambda}\left(a / \rho_{K}(x)\right)^{\lambda} d x=a^{-n} \alpha_{3} V(K)
$$

where

$$
\alpha_{3}=n \int_{0}^{\infty} s^{n-1} p_{\lambda}(s)^{\lambda} d s= \begin{cases}\frac{n}{p} B\left(\frac{n}{p}, \frac{\lambda}{1-\lambda}-\frac{n}{p}\right), & \lambda<1  \tag{2.9}\\ \frac{n}{p} \Gamma\left(\frac{n}{p}\right), & \lambda=1 \\ \frac{n}{p} B\left(\frac{n}{p}, \frac{\lambda}{\lambda-1}+1\right), & \lambda>1\end{cases}
$$

The proof is similar to that of Lemma 2.1.
3. Constrained maximum Rényi entropy. The following lemma presents the solution to the problem of maximizing the $\lambda$-Rényi entropy when the value of the dual mixed volume of a random vector is fixed.

Lemma 3.1. Suppose $K$ is a star body in $\mathbb{R}^{n}$ and real $p, \lambda, c>0$, with $\lambda>\frac{n}{n+p}$. Consider the problem of finding

$$
\max \operatorname{Ent}_{\lambda}(X)
$$

subject to the constraint that $X$ be a random vector in $\mathbb{R}^{n}$, with finite pth moment, such that

$$
\tilde{V}_{-p}(X, K)=c
$$

Then, the unique maximum is achieved by the random vector whose density function is a.e.

$$
b p_{\lambda}\left(a / \rho_{K}\right)
$$

where $b>0$ is chosen so that $b p_{\lambda}\left(a / \rho_{K}\right)$ is a probability density and $a>0$ is chosen so that

$$
\tilde{V}_{-p}\left(b p_{\lambda}\left(a / \rho_{K}\right), K\right)=c
$$

Proof. Suppose $f$ is a probability density on $\mathbb{R}^{n}$ such that $\tilde{V}_{-p}(f, K)=c$. For the sake of notational simplicity let $M$ denote

$$
M=\frac{n}{n+p} \tilde{V}_{-p}(f, K)=\int_{\mathbb{R}^{n}} \rho_{K}^{-p} f d x
$$

and let

$$
g_{\lambda, p}=b p_{\lambda}\left(a / \rho_{K}\right)
$$

where $b>0$ is chosen so that $b p_{\lambda}\left(a / \rho_{K}\right)$ is a probability density and $a>0$ is chosen so that

$$
\begin{equation*}
M=b \int_{\mathbb{R}^{n}} \rho_{K}^{-p} p_{\lambda}\left(a / \rho_{K}\right) d x=\int_{\mathbb{R}^{n}} \rho_{K}^{-p} g_{\lambda, p} d x=\int_{\mathbb{R}^{n}} \rho_{K}^{-p} f d x \tag{3.1}
\end{equation*}
$$

(I) Case $\lambda=1$. From the fact that $g_{1, p}=b e^{-a^{p} / \rho_{K}^{p}}$, and the fact that $f$ and $g_{1, p}$ are probability densities, together with the last identity in (3.1), and the definition of $E n t_{1}$, we have

$$
\begin{aligned}
\int_{\mathbb{R}^{n}} f \log \left(g_{1, p} / f\right) d x & =-\int_{\mathbb{R}^{n}} f \log f d x+\int_{\mathbb{R}^{n}} f \log g_{1, p} d x \\
& =-\int_{\mathbb{R}^{n}} f \log f d x+\int_{\mathbb{R}^{n}}\left(\log b-a^{p} \rho_{K}^{-p}\right) f d x \\
& =-\int_{\mathbb{R}^{n}} f \log f d x+\int_{\mathbb{R}^{n}}\left(\log b-a^{p} \rho_{K}^{-p}\right) g_{1, p} d x \\
& =\operatorname{Ent}_{1}(f)-\operatorname{Ent}_{1}\left(g_{1, p}\right)
\end{aligned}
$$

From the strict concavity of the log function, we see that

$$
\int_{\mathbb{R}^{n}} f \log \left(g_{1, p} / f\right) d x \leq \int_{\mathbb{R}^{n}} \log g_{1, p} d x \leq \log \int_{\mathbb{R}^{n}} g_{1, p} d x=0
$$

with equality if and only if a.e. $f=g_{1, p}$. Thus,

$$
\operatorname{Ent}_{1}(f) \leq \operatorname{Ent}_{1}\left(g_{1, p}\right),
$$

with equality if and only if a.e. $f=g_{1, p}$.
(II) Case $\lambda \neq 1$. From the fact that $g_{\lambda, p}$ is a density function, Lemmas 2.2, 2.3 and 2.1, we have

$$
\begin{aligned}
1 & =\int_{\mathbb{R}^{n}} g_{\lambda, p} d x=b a^{-n} \alpha_{2} V(K), \\
\int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda} d x & =b^{\lambda} a^{-n} \alpha_{3} V(K), \\
M & =\int_{\mathbb{R}^{n}} \rho_{K}^{-p} g_{\lambda, p} d x=b a^{-(n+p)} \alpha_{1} V(K) .
\end{aligned}
$$

Thus,

$$
\begin{equation*}
b^{1-\lambda} \int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda} d x=\frac{\alpha_{3}}{\alpha_{2}} \quad \text { and } \quad a^{p} M=\frac{\alpha_{1}}{\alpha_{2}} \tag{3.2}
\end{equation*}
$$

We now divide the case $\lambda \neq 1$ into two subcases, subcase $\lambda<1$ and subcase $\lambda>1$.

Subcase $\lambda<1$. The Hölder inequality and the fact that $f$ and $g_{\lambda, p}$ are probability densities show that

$$
\begin{equation*}
\int_{\mathbb{R}^{n}} g_{\lambda, p}^{1-\lambda} f^{\lambda} d x \leq\left(\int_{\mathbb{R}^{n}} g_{\lambda, p} d x\right)^{1-\lambda}\left(\int_{\mathbb{R}^{n}} f d x\right)^{\lambda}=1 \tag{3.3}
\end{equation*}
$$

with equality if and only if a.e. $f=g_{\lambda, p}$.
From the definition of $g_{\lambda, p}$ and the definition of $p_{\lambda}$, for $\lambda<1$, we see that $g_{\lambda, p}^{1-\lambda}=b^{1-\lambda}-a^{p} \rho_{K}^{-p} g_{\lambda, p}^{1-\lambda}$. From this and the Hölder inequality, again, we have

$$
\begin{aligned}
\int_{\mathbb{R}^{n}} g_{\lambda, p}^{1-\lambda} f^{\lambda} d x & =b^{1-\lambda} \int_{\mathbb{R}^{n}} f^{\lambda} d x-\int_{\mathbb{R}^{n}} a^{p} \rho_{K}^{-p} g_{\lambda, p}^{1-\lambda} f^{\lambda} d x \\
& \geq b^{1-\lambda} \int_{\mathbb{R}^{n}} f^{\lambda} d x-a^{p}\left(\int_{\mathbb{R}^{n}} \rho_{K}^{-p} g_{\lambda, p} d x\right)^{1-\lambda}\left(\int_{\mathbb{R}^{n}} \rho_{K}^{-p} f d x\right)^{\lambda} \\
& =b^{1-\lambda} \int_{\mathbb{R}^{n}} f^{\lambda} d x-a^{p} M
\end{aligned}
$$

This together with (3.3) shows that

$$
\begin{equation*}
b^{1-\lambda} \int_{\mathbb{R}^{n}} f^{\lambda} d x \leq a^{p} M+1 \tag{3.4}
\end{equation*}
$$

with equality if and only if a.e. $f=g_{\lambda, p}$.
In this subcase (3.2), together with (2.7), (2.8) and (2.9), give

$$
\begin{aligned}
b^{1-\lambda} \int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda} d x & =\frac{\alpha_{3}}{\alpha_{2}}
\end{aligned}=\frac{\lambda /(1-\lambda)}{\lambda /(1-\lambda)-n / p},
$$

It follows that

$$
a^{p} M+1=b^{1-\lambda} \int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda} d x
$$

This and (3.4) give the desired result that

$$
\int_{\mathbb{R}^{n}} f^{\lambda} d x \leq \int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda} d x
$$

with equality if and only if a.e. $f=g_{\lambda, p}$.
Subcase $\lambda>1$. From the Hölder inequality we have

$$
\begin{equation*}
\int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda-1} f d x \leq\left(\int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda} d x\right)^{1-1 / \lambda}\left(\int_{\mathbb{R}^{n}} f^{\lambda} d x\right)^{1 / \lambda} \tag{3.5}
\end{equation*}
$$

with equality if and only if a.e. $f=g_{\lambda, p}$.
From the definition of $g_{\lambda, p}$ and the definition of $p_{\lambda}$, for $\lambda>1$, we see that
in this case $g_{\lambda, p}^{\lambda-1} \geq b^{\lambda-1}\left(1-a^{p} \rho_{K}^{-p}\right)$. This and the fact that $f$ is a probability density give

$$
\begin{equation*}
\int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda-1} f d x \geq b^{\lambda-1}\left(1-a^{p} \int_{\mathbb{R}^{n}} \rho_{K}^{-p} f d x\right)=b^{\lambda-1}\left(1-a^{p} M\right) \tag{3.6}
\end{equation*}
$$

In this subcase (3.2), together with (2.7), (2.8) and (2.9), give

$$
\begin{aligned}
b^{1-\lambda} \int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda} d x & =\frac{\alpha_{3}}{\alpha_{2}}
\end{aligned}=\frac{\lambda /(\lambda-1)}{\lambda /(\lambda-1)+n / p},
$$

These identities yield

$$
b^{\lambda-1}\left(1-a^{p} M\right)=\int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda} d x
$$

and, thus, from (3.6) we have

$$
\int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda-1} f d x \geq \int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda} d x
$$

This and (3.5) give the desired result,

$$
\int_{\mathbb{R}^{n}} f^{\lambda} d x \geq \int_{\mathbb{R}^{n}} g^{\lambda} d x
$$

with equality if and only if a.e. $f=g_{\lambda, p}$.

## 4. Inequalities between dual mixed volumes and Rényi entropy.

Lemma 4.1. Suppose real $p>0$ and $\lambda>\frac{n}{n+p}$. If $K$ is a star body in $\mathbb{R}^{n}$ and $X$ is a random vector in $\mathbb{R}^{n}$ that has finite pth moment, then

$$
\tilde{V}_{-p}(X, K) \geq\left(1+\frac{p}{n}\right) c_{0}\left[N_{\lambda}(X) / V(K)\right]^{p / n}
$$

with equality if and only if $X$ is a random vector with density function a.e. proportional to $p_{\lambda}\left(a / \rho_{K}\right)$, with some $a>0$.

Proof. Abbreviate

$$
M=\int_{\mathbb{R}^{n}} \rho_{K}^{-p} f d x=\frac{n}{n+p} \tilde{V}_{-p}(f, K)
$$

Let

$$
g_{\lambda, p}=b p_{\lambda}\left(a / \rho_{K}\right)
$$

where $b$ is chosen so that $b p_{\lambda}\left(a / \rho_{K}\right)$ is a probability density and $a$ is chosen so that

$$
\begin{equation*}
M=b \int_{\mathbb{R}^{n}} \rho_{K}^{-p} p_{\lambda}\left(a / \rho_{K}\right) d x \tag{4.1}
\end{equation*}
$$

First note that from Lemma 3.1 we know that

$$
\begin{equation*}
N_{\lambda}\left(g_{\lambda, p}\right) \geq N_{\lambda}(f) \tag{4.2}
\end{equation*}
$$

with equality if and only if a.e. $g_{\lambda, p}=f$.
From (4.1) and Lemma 2.1, we see that

$$
M=b a^{-(n+p)} \alpha_{1} V(K) .
$$

From the fact that $g_{\lambda, p}$ is a density function and and the fact that $g_{\lambda, p}=b p_{\lambda}\left(a / \rho_{K}\right)$, together with Lemma 2.2, we see that

$$
1=\int_{\mathbb{R}^{n}} g_{\lambda, p} d x=b a^{-n} \alpha_{2} V(K)
$$

Thus, we have

$$
\begin{align*}
a^{p} & =\frac{\alpha_{1}}{\alpha_{2} M}  \tag{4.3}\\
b & =\frac{a^{n}}{\alpha_{2} V(K)} . \tag{4.4}
\end{align*}
$$

From Lemma 2.3, together with (4.3) and (4.4), we have

$$
\int_{\mathbb{R}^{n}} g_{\lambda, p}^{\lambda} d x=b^{\lambda} a^{-n} \alpha_{3} V(K)=\frac{1}{\alpha_{2}^{\lambda}}\left(\frac{\alpha_{1}}{\alpha_{2} M}\right)^{n(\lambda-1) / p} \alpha_{3} V(K)^{1-\lambda}
$$

Thus, for $\lambda \neq 1$,

$$
\begin{equation*}
N_{\lambda}\left(g_{\lambda, p}\right)^{p / n}=\left(\frac{\alpha_{3}}{\alpha_{2}^{\lambda}}\right)^{p /(n(1-\lambda))} \frac{\alpha_{2}}{\alpha_{1}} M V(K)^{p / n} \tag{4.5}
\end{equation*}
$$

Suppose $\lambda=1$. From the definitions of $g_{1, p}, p_{1}$, Ent ${ }_{1}$ and the fact that $b p_{1}\left(a / \rho_{K}\right)$ is a probability density, together with Lemma 2.1, (4.4) and finally, (2.7) together with (2.8), we have

$$
\begin{aligned}
\operatorname{Ent}_{1}\left(g_{1, p}\right) & =-\int_{\mathbb{R}^{n}} b e^{-a^{p} \rho_{K}^{-p}}\left(\log b-a^{p} \rho_{K}^{-p}\right) d x \\
& =-\log b \int_{\mathbb{R}^{n}} b p_{1}\left(a / \rho_{K}\right)+a^{p} b \int_{\mathbb{R}^{n}} \rho_{K}^{-p} p_{1}\left(a / \rho_{K}\right) d x \\
& =-\log b+b a^{-n} \alpha_{1} V(K) \\
& =-\log b+\frac{\alpha_{1}}{\alpha_{2}} \\
& =-\log b+\frac{n}{p}
\end{aligned}
$$

This and the definition of $N_{1}$, together with (4.3) and (4.4), and finally, (2.7) and (2.8), give

$$
\begin{equation*}
N_{1}\left(g_{1, p}\right)^{p / n}=e b^{-p / n}=\frac{e p}{n} \alpha_{2}^{p / n} M V(K)^{p / n} . \tag{4.6}
\end{equation*}
$$

Therefore, from (4.2) and (4.5), [or from (4.2) and (4.6) when $\lambda=1$ ], we have

$$
M^{1 / p} \geq c_{0}^{\prime}\left[N_{\lambda}(f) / V(K)\right]^{1 / n}
$$

with equality if and only if $f=g_{\lambda, p}$, a.e., where $c_{0}^{\prime}$ is given by

$$
c_{0}^{\prime}= \begin{cases}\left(\frac{\alpha_{3}}{\alpha_{2}}\right)^{1 /(n(\lambda-1))}\left(\frac{\alpha_{1}}{\alpha_{2}}\right)^{1 / p} \alpha_{2}^{-1 / n}, & \lambda \neq 1 \\ \left(\frac{n}{e p}\right)^{1 / p} \alpha_{2}^{-1 / n}, & \lambda=1\end{cases}
$$

By using (2.7)-(2.9) one now obtains the inequality of the lemma.
A simple limit argument shows that the inequality of Lemma 4.1 holds when $\lambda=\infty$. However, in order to obtain the equality conditions we shall proceed in a different manner.

LEMMA 4.2. Suppose $p>0$. If $K$ is a star body in $\mathbb{R}^{n}$ and $X$ is a random vector in $\mathbb{R}^{n}$ that has finite pth moment and bounded density, then

$$
\tilde{V}_{-p}(X, K)^{1 / p} \geq\left[N_{\infty}(X) / V(K)\right]^{1 / n}
$$

with equality if and only if there exists an $a>0$ such that the density function of $X$ is, a.e., $\mathbf{1}_{a K} / V(a K)$.

Proof. Let $f$ be the probability density of $X$, and let $\|f\|_{\infty}$ denote the essential supremum of $f$. We are assuming that $\|f\|_{\infty}<\infty$, and for convenience let $a=\left[\|f\|_{\infty} V(K)\right]^{-1 / n}$.

From the definition of a radial function (and the fact that $K$ is a star body) it follows immediately that $x \in \operatorname{int} a K$ if and only if $a>\rho_{K}^{-1}(x)$ or equivalently int $a K$ is precisely the set on which the function $\left(a^{p}-\rho_{K}^{-p}\right)_{+}$is positive. This observation and the fact that $f$ is a density function, together with definition (2.6), show that

$$
\int_{\mathbb{R}^{n}}\left(a^{p}-\rho_{K}^{-p}\right)_{+} f d x \geq \int_{\mathbb{R}^{n}}\left(a^{p}-\rho_{K}^{-p}\right) f d x=a^{p}-\frac{n}{n+p} \tilde{V}_{-p}(f, K),
$$

with equality if and only if $f(x)=0$ for almost all $x \notin$ int $a K$.
Again, from the fact that int $a K$ is precisely the set on which the function $\left(a^{p}-\rho_{K}^{-p}\right)_{+}$is positive, together with (2.2), the fact that $\tilde{V}_{-p}(\cdot, K)$ is homogeneous of degree $n+p$, and (2.3) we have

$$
\begin{aligned}
\int_{\mathbb{R}^{n}}\left(a^{p}-\rho_{K}^{-p}\right)_{+} f d x & \leq\|f\|_{\infty} \int_{\mathbb{R}^{n}}\left(a^{p}-\rho_{K}^{-p}\right)_{+} d x \\
& =\|f\|_{\infty} \int_{a K}\left(a^{p}-\rho_{K}^{-p}\right) d x
\end{aligned}
$$

$$
\begin{aligned}
& =\|f\|_{\infty} a^{p} \int_{a K} 1 d x-\frac{n\|f\|_{\infty}}{n+p} \tilde{V}_{-p}(a K, K) \\
& =\frac{p}{n+p} a^{n+p}\|f\|_{\infty} V(K),
\end{aligned}
$$

with equality if and only if $f$ is a.e. constant on int $a K$.
Combining the above inequalities and recalling that $a=\left[\|f\|_{\infty} V(K)\right]^{-1 / n}$, give

$$
\tilde{V}_{-p}(f, K) \geq \frac{n+p}{n} a^{p}-\frac{p}{n} a^{n+p}\|f\|_{\infty} V(K)=\left[\|f\|_{\infty} V(K)\right]^{-p / n},
$$

with equality if and only if $f$ is a.e. constant on int $a K$ and 0 on its complement.
This and the fact that $N_{\infty}(f)=1 /\|f\|_{\infty}$ give the desired inequality and the fact that equality holds if and only if, a.e., $f=\mathbf{1}_{a K} / V(a K)$.
5. A star body associated with a random vector. Suppose $p>0$, and $f: \mathbb{R}^{n} \rightarrow \mathbb{R}$ is a density function that has finite $p$ th moment. Define the Borel measure $\mu_{f}$ on $S^{n-1}$ by letting

$$
\int_{S^{n-1}} q(u) d \mu_{f}(u)=\int_{\mathbb{R}^{n}} f(x) q(x /|x|)|x|^{p} d x
$$

for each $q \in C\left(S^{n-1}\right)$. Since the measure $\mu_{f}$ is absolutely continuous with respect to spherical Lebesgue measure, there is an essentially unique function $\bar{f} \in L_{1}\left(S^{n-1}\right)$ such that $\bar{f} \geq 0$ and

$$
\frac{1}{n} \int_{S^{n-1}} q(u) \bar{f}(u) d u=\int_{\mathbb{R}^{n}} f(x) q(x /|x|)|x|^{p} d x
$$

for each $q \in C\left(S^{n-1}\right)$. Thus, from $f$ we get a unique $L_{n+p}$-star, $T_{p} f$, defined by $\rho_{T_{p} f}^{n+p}=\bar{f}$ such that

$$
\begin{equation*}
\tilde{V}_{-p}\left(T_{p} f, Q\right)=\int_{\mathbb{R}^{n}} f(x) \rho_{Q}(x)^{-p} d x \tag{5.1}
\end{equation*}
$$

for each star body $Q$. Note that $0<V\left(T_{p} f\right)<\infty$ [since $V\left(T_{p} f\right)=\frac{1}{n} \times$ $\int_{S^{n-1}} \bar{f}^{n /(n+p)}$ ], and define $S_{p} f=V\left(T_{p} f\right)^{1 / p} T_{p} f$. The homogeneity (of degree $n$ ) of volume now immediately gives $V\left(S_{p} f\right)^{1 /(n+p)}=V\left(T_{p} f\right)^{1 / p}$. But $V\left(S_{p} f\right)^{-1 /(n+p)} S_{p} f=T_{p} f$ and the fact that $\tilde{V}_{-p}(\cdot, Q)$ is homogeneous of degree $n+p$, let us rewrite (5.1) as

$$
\begin{equation*}
\tilde{V}_{-p}\left(S_{p} f, Q\right) / V\left(S_{p} f\right)=\int_{\mathbb{R}^{n}} f(x) \rho_{Q}(x)^{-p} d x \tag{5.2}
\end{equation*}
$$

or by (2.2) equivalently,

$$
\begin{equation*}
\frac{n+p}{n V\left(S_{p} f\right)} \int_{S_{p} f} \rho_{Q}(x)^{-p} d x=\int_{\mathbb{R}^{n}} f(x) \rho_{Q}(x)^{-p} d x \tag{5.3}
\end{equation*}
$$

for each star body $Q$.
If $X$ is a random vector in $\mathbb{R}^{n}$ with density function $f$, that has finite $p$ th moment, then we will often write $S_{p} X$ rather than $S_{p} f$. Thus, from (5.2) and (2.6) we see that $S_{p} X$ can be defined by simply requiring that

$$
\begin{equation*}
\tilde{V}_{-p}\left(S_{p} X, Q\right) / V\left(S_{p} X\right)=\left(1+\frac{p}{n}\right)^{-1} \tilde{V}_{-p}(X, Q) \tag{5.4}
\end{equation*}
$$

hold for each star body $Q$.
Observe that if $f \in C_{c}^{\infty}\left(\mathbb{R}^{n}\right)$ is positive in a neighborhood of the origin, then the $L_{n+p}$-star, $S_{p} f$, is in fact a star body, and it follows that from (2.2), together with (5.2), its radial function $\rho_{S_{p} X}$ is given by

$$
\begin{equation*}
\frac{1}{V\left(S_{p} X\right)} \rho_{S_{p} X}(u)^{n+p}=n \int_{0}^{\infty} f(r u) r^{n+p-1} d r . \tag{5.5}
\end{equation*}
$$

If $f \in C_{c}^{\infty}\left(\mathbb{R}^{n}\right)$ is positive in a neighborhood of the origin, then by taking $Q=S_{p} X$ in (5.4) and recalling (2.3) we see that

$$
\begin{equation*}
\tilde{V}_{-p}\left(X, S_{p} X\right)=1+\frac{p}{n} \tag{5.6}
\end{equation*}
$$

As will now be shown, the volume of the star associated with a random vector can be bounded from below by the $\lambda$-Rényi entropy power of the random vector. Although the inequality of our next lemma is stated without equality conditions, it is sharp.

Lemma 5.1. Suppose real $p>0$ and real $\lambda>\frac{n}{n+p}$. If $X$ is a random vector in $\mathbb{R}^{n}$ that has finite pth moment, then

$$
V\left(S_{p} X\right) \geq c_{0}^{n / p} N_{\lambda}(X)
$$

Proof. Let $f$ denote the density function of $X$. First note that if $f \in C_{c}^{\infty}\left(\mathbb{R}^{n}\right)$ is positive in a neighborhood of the origin, then from (5.6) and Lemma 4.1,

$$
\begin{equation*}
\left(1+\frac{p}{n}\right)^{n / p}=\tilde{V}_{-p}\left(X, S_{p} X\right)^{n / p} \geq\left(1+\frac{p}{n}\right)^{n / p} c_{0}^{n / p} N_{\lambda}(X) / V\left(S_{p} X\right) \tag{5.7}
\end{equation*}
$$

This establishes the desired inequality for the special case where $f \in C_{c}^{\infty}\left(\mathbb{R}^{n}\right)$ is positive in a neighborhood of the origin.

To handle the case for arbitrary $f$ when $\lambda \neq 1$, choose a sequence of probability density functions $f_{i} \in C_{c}^{\infty}\left(\mathbb{R}^{n}\right)$ that are positive in a neighborhood of the origin and such that

$$
\begin{equation*}
\lim _{i \rightarrow \infty} \int_{\mathbb{R}^{n}} f_{i}(x) \rho_{Q}(x)^{-p} d x=\int_{\mathbb{R}^{n}} f(x) \rho_{Q}(x)^{-p} d x \tag{5.8}
\end{equation*}
$$

for each star body $Q$, and

$$
\begin{equation*}
\limsup _{i \rightarrow \infty} N_{\lambda}\left(f_{i}\right) \geq N_{\lambda}(f) \tag{5.9}
\end{equation*}
$$

Now suppose $Q$ is a star body such that $V(Q)=1$. From (5.8) and (5.2), followed by the dual mixed volume inequality (2.4), then (5.7), and finally (5.9), we have

$$
\begin{aligned}
& \tilde{V}_{-p}\left(S_{p} f, Q\right) / V\left(S_{p} f\right) \\
& \quad=\lim _{i \rightarrow \infty} \tilde{V}_{-p}\left(S_{p} f_{i}, Q\right) / V\left(S_{p} f_{i}\right) \geq \limsup _{i \rightarrow \infty} V\left(S_{p} f_{i}\right)^{p / n} \\
& \quad \geq c_{0} \limsup _{i \rightarrow \infty} N_{\lambda}\left(f_{i}\right)^{p / n} \geq c_{0} N_{\lambda}(f)^{p / n}
\end{aligned}
$$

This, together with (2.5), now completes the proof for the case of arbitrary $f$ when $\lambda \neq 1$. The case of arbitrary $f$ when $\lambda=1$ now follows from the case $\lambda \neq 1$ by taking a simple limit.

It will be convenient to re-define $c_{0}$ so that it is defined not only for positive $\lambda$ but for $\lambda=\infty$ as well. To this end, define

Taking $\lambda \rightarrow \infty$ in Lemma 5.1 [and noting that $\lim _{\lambda \rightarrow \infty} c_{0}=$ $(1+p / n)^{-1}$ ] gives

Lemma 5.2. Suppose real $p>0$ and $X$ is a random vector in $\mathbb{R}^{n}$ that has finite pth moment, then

$$
V\left(S_{p} X\right) \geq c_{0}^{n / p} N_{\infty}(X)
$$

6. Moments of random variables. Suppose $p \geq 1$. If $K \subset \mathbb{R}^{n}$ is an $L_{n+p}$-star that is of positive volume, then its polar $L_{p}$-centroid body, $\Gamma_{p}^{*} K$, is the convex body whose radial function is given by

$$
\begin{equation*}
\rho_{\Gamma_{p}^{*} K}(u)^{-p}=\frac{1}{V(K)} \int_{K}|u \cdot x|^{p} d x . \tag{6.1}
\end{equation*}
$$

From this definition it is easily seen that if $E$ is an ellipsoid centered at the origin, then

$$
\begin{equation*}
\Gamma_{p}^{*} E \text { is a dilate of } E^{*} \tag{6.2}
\end{equation*}
$$

Definitions (2.2), (6.1) and Fubini's theorem show that for positive-volume $L_{n+p}$-stars $K, L \subset \mathbb{R}^{n}$, we have

$$
\begin{equation*}
\tilde{V}_{-p}\left(K, \Gamma_{p}^{*} L\right) / V(K)=\tilde{V}_{-p}\left(L, \Gamma_{p}^{*} K\right) / V(L) \tag{6.3}
\end{equation*}
$$

Take $L=\Gamma_{p}^{*} K$ in (6.3) and from (2.3) get the following: If $K \subset \mathbb{R}^{n}$ is an $L_{n+p}$-star, then

$$
\begin{equation*}
V(K)=\tilde{V}_{-p}\left(K, \Gamma_{p}^{*} \Gamma_{p}^{*} K\right) \tag{6.4}
\end{equation*}
$$

For $p=1$ and a more restricted class of bodies, identity (6.3) was first given in Lutwak (1990).

Lemma 6.1. Let $X$ and $Y$ be independent random variables in $\mathbb{R}^{n}$ that have finite pth moment, then for $p \geq 1$,

$$
\tilde{V}_{-p}\left(X, \Gamma_{p}^{*} S_{p} Y\right)=E\left(|X \cdot Y|^{p}\right)=\tilde{V}_{-p}\left(Y, \Gamma_{p}^{*} S_{p} X\right)
$$

Proof. Let $f, g$ be the density functions of $X, Y$.
First note that an obvious limit argument in (5.3) shows that for each $x \in \mathbb{R}^{n}$,

$$
\begin{equation*}
\frac{n+p}{n V\left(S_{p} g\right)} \int_{S_{p} g}|x \cdot y|^{p} d y=\int_{\mathbb{R}^{n}} g(y)|x \cdot y|^{p} d y \tag{6.5}
\end{equation*}
$$

By using (2.6), (6.1) and (6.5), we have

$$
\begin{aligned}
\tilde{V}_{-p}\left(X, \Gamma_{p}^{*} S_{p} Y\right) & =\left(1+\frac{p}{n}\right) \int_{\mathbb{R}^{n}} \rho_{\Gamma_{p}^{*} S_{p} Y}^{-p}(x) f(x) d x \\
& =\frac{n+p}{n V\left(S_{p} Y\right)} \int_{\mathbb{R}^{n}} f(x) \int_{S_{p} Y}|x \cdot y|^{p} d y d x \\
& =\int_{\mathbb{R}^{n}} \int_{\mathbb{R}^{n}}|x \cdot y|^{p} f(x) g(y) d x d y=E\left(|X \cdot Y|^{p}\right)
\end{aligned}
$$

To complete the proof observe that $E\left(|X \cdot Y|^{p}\right)$ is symmetric in $X$ and $Y$.
Define $c_{2}$ by

$$
c_{2}=\left(\frac{\pi^{1 / 2+p}(1+p / n) \Gamma\left(\frac{p+n}{2}\right)}{\Gamma\left(\frac{n}{2}\right) \Gamma\left(\frac{p+1}{2}\right) \Gamma\left(1+\frac{n}{2}\right)^{2 p / n}}\right)^{n / p} .
$$

We shall use the following result of Lutwak and Zhang (1997) [see also Lutwak, Yang and Zhang (2000), as well as Campi and Gronchi (2002)]: If $K$ is an originsymmetric convex body, then

$$
\begin{equation*}
V\left(\Gamma_{p}^{*} K\right) V(K) \leq c_{2}, \tag{6.6}
\end{equation*}
$$

with equality if and only if $K$ is an ellipsoid.

THEOREM 6.2. Suppose real $p \geq 1$ and $\lambda \in\left(\frac{n}{n+p}, \infty\right]$. Let $X$ and $Y$ be independent random variables in $\mathbb{R}^{n}$ that have finite pth moment (and are also of bounded density if $\lambda=\infty$ ), then

$$
E\left(|X \cdot Y|^{p}\right) \geq c_{0}^{2} c_{2}^{-p / n}\left(1+\frac{p}{n}\right)\left[N_{\lambda}(X) N_{\lambda}(Y)\right]^{p / n}
$$

with equality for $\lambda<\infty$ if and only if there exists a $\phi \in G L(n)$ with $X$ having density a.e. $\phi_{p, \lambda}$ and $Y$ having density a.e. $\left(a \phi^{-t}\right)_{p, \lambda}$, with $a>0$, and equality for $\lambda=\infty$ if and only if the densities of $X$ and $Y$ are a.e. proportional to the characteristic functions of dilates of polar-reciprocal origin-centered ellipsoids.

Proof. With $i \neq j$, let $\left\{X_{i}, X_{j}\right\}=\{X, Y\}$ and let $f_{i}, f_{j}$ denote the density functions of $X_{i}, X_{j}$. From Lemma 6.1, (5.4), the dual mixed volume inequality (2.4), Lemma 5.1, (6.6), (2.4), (6.4) and Lemma 5.1, we have

$$
\begin{aligned}
E\left(|X \cdot Y|^{p}\right)= & \tilde{V}_{-p}\left(X_{i}, \Gamma_{p}^{*} S_{p} X_{j}\right) \\
= & \left(1+\frac{p}{n}\right) \tilde{V}_{-p}\left(S_{p} X_{i}, \Gamma_{p}^{*} S_{p} X_{j}\right) / V\left(S_{p} X_{i}\right) \\
\geq & \left(1+\frac{p}{n}\right)\left[V\left(S_{p} X_{i}\right) / V\left(\Gamma_{p}^{*} S_{p} X_{j}\right)\right]^{p / n} \\
\geq & \left(1+\frac{p}{n}\right) c_{0}\left[N_{\lambda}\left(X_{i}\right) / V\left(\Gamma_{p}^{*} S_{p} X_{j}\right)\right]^{p / n} \\
\geq & \left(1+\frac{p}{n}\right) c_{0} c_{2}^{-p / n} N_{\lambda}\left(X_{i}\right)^{p / n} V\left(\Gamma_{p}^{*} \Gamma_{p}^{*} S_{p} X_{j}\right)^{p / n} \\
\geq & \left(1+\frac{p}{n}\right) c_{0} c_{2}^{-p / n} N_{\lambda}\left(X_{i}\right)^{p / n} \tilde{V}_{-p}\left(S_{p} X_{j}, \Gamma_{p}^{*} \Gamma_{p}^{*} S_{p} X_{j}\right)^{-1} \\
& \times V\left(S_{p} X_{j}\right)^{(n+p) / n} \\
= & \left(1+\frac{p}{n}\right) c_{0} c_{2}^{-p / n} N_{\lambda}\left(X_{i}\right)^{p / n} V\left(S_{p} X_{j}\right)^{p / n} \\
\geq & \left(1+\frac{p}{n}\right) c_{0}^{2} p c_{2}^{-p / n} N_{\lambda}\left(X_{i}\right)^{p / n} N_{\lambda}\left(X_{j}\right)^{p / n} .
\end{aligned}
$$

Suppose there is equality in the inequality of the theorem. Obviously, this implies equality in all of the inequalities above. Equality in the first inequality, by the equality conditions of the dual Minkowski inequality (2.4), implies that there exist $d_{i}>0$ such that, a.e.,

$$
\begin{equation*}
\rho_{S_{p} X_{i}}=d_{i} \rho_{\Gamma_{p}^{*} S_{p} X_{j}} \tag{6.7}
\end{equation*}
$$

The fact that $\tilde{V}_{-p}\left(X_{i}, \Gamma_{p}^{*} S_{p} X_{j}\right)=\left(1+\frac{p}{n}\right) c_{0}\left[N_{\lambda}\left(X_{i}\right) / V\left(\Gamma_{p}^{*} S_{p} X_{j}\right)\right]^{p / n}$ and the equality conditions of Lemma 4.1 (or Lemma 4.2 if $\lambda=\infty$ ) show that there
exist $a_{i}, b_{i}>0$ such that, a.e.,

$$
f_{i}= \begin{cases}b_{i} p_{\lambda}\left(a_{i} / \rho_{\Gamma_{p}^{*} S_{p} X_{j}}\right), & \lambda<\infty  \tag{6.8}\\ \mathbf{1}_{a_{i} \Gamma_{p}^{*} S_{p} X_{j}} / V\left(a_{i} \Gamma_{p}^{*} S_{p} X_{j}\right), & \lambda=\infty .\end{cases}
$$

From the equality conditions of (6.6), we see that equality in the third inequality implies that there exits an origin-centered ellipsoid, $E_{j}$, such that

$$
\begin{equation*}
\Gamma_{p}^{*} S_{p} X_{j}=E_{j} \tag{6.9}
\end{equation*}
$$

But (6.9) and (6.8) show that, a.e.,

$$
f_{i}= \begin{cases}b_{i} p_{\lambda}\left(a_{i} / \rho_{E_{j}}\right), & \lambda<\infty  \tag{6.10}\\ \mathbf{1}_{a_{i} E_{j}} / V\left(a_{i} E_{j}\right), & \lambda=\infty\end{cases}
$$

But (6.7) and (6.9) show that $\rho_{S_{p} X_{i}}=d_{i} \rho_{E_{j}}$, a.e., and, hence, by (6.1) and (6.2)

$$
\begin{equation*}
\Gamma_{p}^{*} S_{p} X_{i}=d_{i}^{\prime} E_{j}^{*} \tag{6.11}
\end{equation*}
$$

for some $d_{i}^{\prime}>0$.
Note that (6.8) shows that, a.e.,

$$
f_{j}= \begin{cases}b_{j} p_{\lambda}\left(a_{j} / \rho_{\Gamma_{p}^{*} S_{p} X_{i}}\right), & \lambda<\infty \\ \mathbf{1}_{a_{j} \Gamma_{p}^{*} S_{p} X_{i}} / V\left(a_{j} \Gamma_{p}^{*} S_{p} X_{i}\right), & \lambda=\infty\end{cases}
$$

and when combined with (6.11) we see that, a.e.,

$$
f_{j}= \begin{cases}b_{j} p_{\lambda}\left(a_{j} / \rho_{d_{i}^{\prime}} E_{j}^{*}\right), & \lambda<\infty,  \tag{6.12}\\ \mathbf{1}_{a_{j}^{\prime} E_{j}^{*}} / V\left(a_{j}^{\prime} E_{j}^{*}\right), & \lambda=\infty,\end{cases}
$$

for some $a_{j}^{\prime}>0$. When (6.10) and (6.12) are combined with (2.1), we get the desired equality conditions.

A comment follows regarding the above proof. Two extra steps were needed in the proof of the inequality of Theorem 6.2 because inequality (6.6) had not been established for the class of $L_{n+p}$-stars. The class reduction technique that is used to overcome this difficulty was introduced in Lutwak (1986).

Suppose $K, L \subset \mathbb{R}^{n}$ are compact. In Theorem 6.2, with $\lambda=\infty$, let $X$ and $Y$ be random vectors with density $\mathbf{1}_{K} / V(K)$ and $\mathbf{1}_{L} / V(L)$ and get

COROLLARY 6.3. If $K$ and $L$ are compact sets in $\mathbb{R}^{n}$, then for real $p \geq 1$,

$$
\int_{K} \int_{L}|x \cdot y|^{p} d x d y \geq \frac{n}{n+p} c_{2}^{-p / n}[V(K) V(L)]^{(n+p) / n}
$$

with equality if and only if $K$ and $L$ are, up to sets of measure 0 , dilates of polarreciprocal, origin-centered ellipsoids.

The limiting case $p \rightarrow \infty$ of Corollary 6.3 gives
Corollary 6.4. If $K$ and $L$ are compact sets in $\mathbb{R}^{n}$, then

$$
\max _{x \in K, y \in L}|x \cdot y| \geq \omega_{n}^{-2 / n}[V(K) V(L)]^{1 / n}
$$

If $K$ is an origin-symmetric convex body and $L$ is its polar, then the above inequality is the classical Blaschke-Santaló inequality (and $\omega_{n}^{-2 / n}$ is the precise constant that makes the inequality sharp).

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